

HORIZONTAL GEODESICS IN LORENTZIAN HEISENBERG GROUP Heis^3

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ABSTRACT. In this paper, we study different geodesic lines on the Lorentzian Heis^3 . We consider the Lorentzian left invariant metric and use some results of Levi-Civita connection and curvature tensor to present solution of equations for geodesic lines in Lorentzian Heis^3 . Furthermore it is obtained equations of geodesic lines with respect to the left invariant Lorentzian metric of the Heis^3 . We characterize horizontal geodesic curves in Lorentzian Heis^3 and prove that there exist no null horizontal geodesic curves in Lorentzian Heis^3 .

1. INTRODUCTION

Rahmani (see [4]) described one motivation for our study that there are three classes of left invariant Lorentzian metrics on the Heis^3 .

In (see [2]) V. Marenich used Heisenberg left invariant metric and obtained geodesic lines in Heis^3 .

Returning to geometry, we note that the Heis^3 serves also as a contact manifold. This is the odd dimension analog for a symplectic structure. The 1-form $\omega^1 = dt + xdy$, which annihilates e_1 and e_2 is a contact form, i.e. $\omega^1 \wedge d\omega^1$ never vanishes. The 2-form $\Omega = d\omega^1 = dx \wedge dy$ is a symplectic form on the distribution generated by e_1 and e_2 . It can be considered also as a magnetic field. Then the Maxwell's equation can be written as $d\Omega = 0$.

In this paper, we study different geodesic lines on the Lorentzian Heis^3 . We consider the Lorentzian left invariant metric and use some results of Levi-Civita connection and curvature tensor to present solution of equations for geodesic lines in Lorentzian Heis^3 . Furthermore it is obtained equations of geodesic lines with respect to the left invariant Lorentzian metric of the Heis^3 . We characterize horizontal geodesic curves in Lorentzian Heis^3 and prove that there exist no null horizontal geodesic curves in Lorentzian Heis^3 .

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2. Lorentzian Heisenberg Group $\overline{\text{Heis}}^3$

This group is formed by all matrices of the form

$$\text{Heis}^3 = \left\{ \begin{pmatrix} 1 & x & t \\ 0 & 1 & y \\ 0 & 0 & 1 \end{pmatrix} : x, y, t \in \mathbb{R} \right\}$$

with the group multiplication induced by the standard matrix product.

The Lorentzian $\overline{\text{Heis}}^3$ can be seen as the space \mathbb{R}^3 endowed with multiplication

$$(\bar{x}, \bar{y}, \bar{t})(x, y, t) = (\bar{x} + x, \bar{y} + y, \bar{t} + t - \bar{x}y + x\bar{y}).$$

The orthonormal basis

$$E_1 = \begin{pmatrix} 0 & 1 & 0 \\ 0 & 0 & 0 \\ 0 & 0 & 0 \end{pmatrix}, \quad E_2 = \begin{pmatrix} 0 & 0 & 0 \\ 0 & 0 & 1 \\ 0 & 0 & 0 \end{pmatrix}, \quad E_3 = \begin{pmatrix} 0 & 0 & 1 \\ 0 & 0 & 0 \\ 0 & 0 & 0 \end{pmatrix}$$

of tangent space at the identity, determines on $\overline{\text{Heis}}^3$ a left invariant Lorentzian metric

$$(2.1) \quad ds^2 = (\omega^1)^2 + (\omega^2)^2 - (\omega^3)^2$$

where

$$\omega^1 = dt + xdy, \quad \omega^2 = dy, \quad \omega^3 = dx$$

is the left invariant orthonormal coframe associated with the orthonormal left invariant frame,

$$(2.2) \quad e_1 = \frac{\partial}{\partial t}, \quad e_2 = \frac{\partial}{\partial y} - x \frac{\partial}{\partial t}, \quad e_3 = \frac{\partial}{\partial x}.$$

Now let C_{ij}^k be the structure's constants of the Lie algebra g of G (see [5]), that is

$$[e_i, e_j] = C_{ij}^k e_k.$$

The corresponding Lie brackets are

$$[e_2, e_3] = 2e_1, \quad [e_1, e_3] = [e_1, e_2] = 0.$$

The Coshul formula for the Levi-Civita connection is:

$$2g(\nabla_{e_i} e_j, e_k) = C_{ij}^k - C_{jk}^i + C_{ki}^j := L_{ij}^k$$

where the non zero L_{ij}^k 's are

$$(2.3) \quad L_{12}^3 = 2, \quad L_{21}^3 = 2, \quad L_{13}^2 = 2, \quad L_{31}^2 = 2, \quad L_{23}^1 = 2, \quad L_{32}^1 = -2.$$

The element zero $0 = (0, 0, 0)$ is the unit of this group structure and the inverse element for (z, t) is $(z, t)^{-1} = (-z, -t)$. Let $a = (z, t)$, $b = (w, s)$. The commutator of the elements $a, b \in \overline{\text{Heis}}^3$ is equal to

$$\begin{aligned} [a, b] &= aba^{-1}b^{-1} \\ &= (z, t)(w, s)(-z, -t)(-w, -s) \\ &= (z + w - z + w, t + s - t - s + \alpha) = (0, \alpha) \end{aligned}$$

where $\alpha \neq 0$ in general. Which shows that $\overline{\text{Heis}}^3$ is not abelian. On the other hand for any $a, b, c \in \overline{\text{Heis}}^3$, their double commutator is

$$[[a, b], c] = (0, 0).$$

Heisenberg group plays an important role in many branches of mathematics such as representation theory, harmonic analysis, PDEs or even quantum mechanic. $Heis^3$ has a rich geometric structure. In fact its group of isometries is of dimension 4, which is the maximal possible dimension for a non-constant curvature metric on a 3-manifold. Also, from the algebraic point of view, $Heis^3$ is a 2-step nilpotent Lie group, i.e. “almost Abelian”.

3. Horizontal Geodesics in the Lorentzian $Heis^3$

Consider a nonintegrable 2-dimensional distribution $x \rightarrow \mathcal{H}_x$ defined as $\mathcal{H} = \ker \omega$, where ω is a 1-form on $Heis^3$. The distribution \mathcal{H} is called the horizontal distribution. A curve $s \rightarrow c(s) = (x(s), y(s), t(s))$ is called horizontal curve if $c'(s) \in \mathcal{H}_{c(s)}$, for every s (see [6]).

Theorem 3.1. *Spacelike geodesic lines issuing from 0 in the Lorentzian $Heis^3$ is solution following differential equation system:*

$$\begin{aligned}
 (3.1) \quad x'(s) &= \sqrt{1 - \xi^2} \sinh(2\xi s + \phi), \\
 y'(s) &= \sqrt{1 - \xi^2} \cosh(2\xi s + \phi), \\
 t'(s) &= \xi - x(s) \sqrt{1 - \xi^2} \cosh(2\xi s + \phi).
 \end{aligned}$$

Proof. Let $c(s)$ be such a geodesics with a natural parameters s and its vector of velocity given by

$$(3.2) \quad c'(s) = \xi(s) e_1 + \eta(s) e_2 + \rho(s) e_3.$$

Then the equation of a geodesic $\nabla_{c'(s)} c'(s) \equiv 0$ and our table of covariant derivatives (3.2) give

$$(3.3) \quad \xi'(s) e_1 + (\eta'(s) + 2\xi(s) \rho(s)) e_2 + (\rho'(s) - 2\xi(s) \eta(s)) e_3 = 0.$$

Thus we easily obtain the following equations for coordinates of the velocity of the geodesic $c(s)$ in our left-invariant moving frame

$$\begin{aligned}
 (3.4) \quad \xi'(s) &= 0, \\
 \eta'(s) + 2\xi(s) \rho(s) &= 0, \\
 \rho'(s) - 2\xi(s) \eta(s) &= 0
 \end{aligned}$$

or

$$\begin{aligned}
 (3.5) \quad (\eta(s) + \rho(s))' + 2\xi(s) (\rho(s) - \eta(s)) &= 0, \\
 (\eta(s) - \rho(s))' + 2\xi(s) (\rho(s) + \eta(s)) &= 0.
 \end{aligned}$$

Since $c : I \rightarrow Heis^3$ is spacelike curve, we have

$$(3.6) \quad \xi^2(s) + \eta^2(s) - \rho^2(s) = 1$$

and from (3.4) we could take

$$\xi = const.$$

where $|\xi| \leq 1$. From (3.6) we could find that,

$$(3.7) \quad \begin{aligned} \xi &= \text{const.} \\ \eta(s) &= r \cosh(2\xi s + \phi), \\ \rho(s) &= r \sinh(2\xi s + \phi) \end{aligned}$$

where $r = \sqrt{\eta^2(s) - \rho^2(s)}$.

To find equations for geodesic $c(s) = (x(s), y(s), z(s))$ issuing from 0, we note that if

$$(3.8) \quad c'(s) = \xi(s) e_1 + \eta(s) e_2 + \rho(s) e_3$$

and our left-invariant vector fields are

$$e_1 = \frac{\partial}{\partial z}, \quad e_2 = \frac{\partial}{\partial y} - x \frac{\partial}{\partial z}, \quad e_3 = \frac{\partial}{\partial x}.$$

Then we easily have (3.1).

Corollary 3.2. *Let $\gamma : I \rightarrow \text{Heis}^3$ be a unit speed spacelike geodesic curve. Then, the parametric equation of spacelike geodesic are*

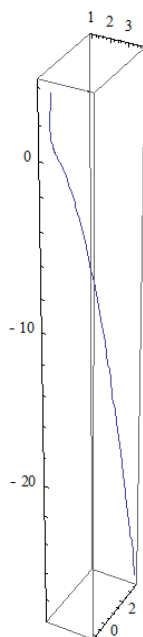
$$(3.9) \quad \begin{aligned} x(s) &= \frac{\sqrt{1-\xi^2}}{2\xi} \cosh(2\xi s + \phi) + c_1, \\ y(s) &= \frac{\sqrt{1-\xi^2}}{2\xi} \sinh(2\xi s + \phi) + c_2, \\ t(s) &= \left(\xi - \frac{1-\xi^2}{4\xi} \right) s - \frac{1-\xi^2}{8\xi^2} \sinh 2(2\xi s + \phi) \\ &\quad - \frac{c_1 \sqrt{1-\xi^2}}{2\xi} \sinh(2\xi s + \phi), \end{aligned}$$

where c_1, c_2, c_3 are constants of integration.

Next, we apply Corollary 3.2.

Example 3.3. *Let us consider parametric equation of spacelike geodesic lines with $\xi = \frac{1}{2}$, $\phi = 0$ and $c_1 = c_2 = c_3 = 1$. Then, we can draw spacelike geodesic*

lines with helping the programme of Mathematica as follow



Theorem 3.4. *If $c : I \rightarrow Heis^3$ spacelike horizontal geodesic lines issuing from zero 0, then satisfy to the following equations:*

$$\begin{aligned}
 (3.10) \quad & x(s) = s \sinh \phi + a, \\
 & y(s) = s \cosh \phi + b, \\
 & t(s) = \frac{s^2}{2} \sinh \phi \cosh \phi + a \cosh \phi s + c,
 \end{aligned}$$

where a, b, c are constants of integration.

Proof. Let $c(s)$ be a horizontal geodesics. From horizontal curve equation, we get

$$(3.11) \quad w(c'(s)) = \xi = 0$$

If we use (3.11) in (3.7) and some calculations, we have

$$\begin{aligned}
 (3.12) \quad & \xi = 0 \\
 & \eta(s) = \cosh \phi, \\
 & \rho(s) = \sinh \phi
 \end{aligned}$$

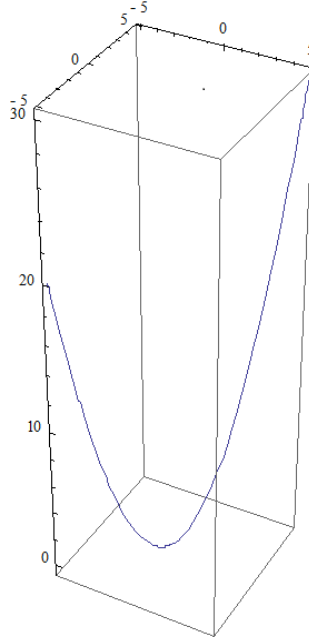
and (3.12) in (3.8) we obtain

$$\begin{aligned}
 & x'(s) = \sinh \phi \\
 & y'(s) = \cosh \phi \\
 & t'(s) = x(s) \cosh \phi.
 \end{aligned}$$

The integration is immediate and yields (3.10).

Next, we apply Theorem 3.4.

Example 3.5. Let us consider parametric equation of spacelike horizontal geodesic lines with $a = b = c = 1$. Then, we can draw spacelike horizontal geodesic with helping the programme of Mathematica as follow



Lemma 3.6. Timelike geodesic lines issuing from 0 in the Lorentzian $Heis^3$ is solution following differential equation system;

$$(3.13) \quad \begin{aligned} x'(s) &= \sqrt{1 + \xi^2} \cosh(2\xi s + \phi), \\ y'(s) &= \sqrt{1 + \xi^2} \sinh(2\xi s + \phi), \\ t'(s) &= \xi - x(s) \sqrt{1 + \xi^2} \sinh(2\xi s + \phi). \end{aligned}$$

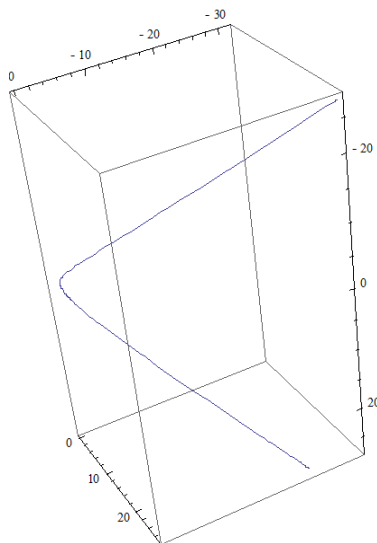
Corollary 3.7. Let $\gamma : I \rightarrow Heis^3$ be a unit speed geodesic curve. Then, the parametric equation of spacelike geodesic are

$$(3.14) \quad \begin{aligned} x(s) &= \frac{\sqrt{1 + \xi^2}}{2\xi} \sinh(2\xi s + \phi) + c_1, \\ y(s) &= \frac{\sqrt{1 + \xi^2}}{2\xi} \cosh(2\xi s + \phi) + c_2, \\ t(s) &= \left(\xi - \frac{1 + \xi^2}{4\xi} \right) s - \frac{1 + \xi^2}{8\xi^2} \sinh 2(2\xi s + \phi) \\ &\quad - \frac{c_1 \sqrt{1 + \xi^2}}{2\xi} \cosh(2\xi s + \phi), \end{aligned}$$

where c_1, c_2, c_3 are constants of integration.

Next, we apply Corollary 3.7.

Example 3.8. Let us consider parametric equation of spacelike geodesic lines with $\xi = 1$, $\phi = 0$ and $c_1 = c_2 = c_3 = 1$. Then, we can draw spacelike geodesic lines with helping the programme of Mathematica as follow



Corollary 3.9. If $c : I \rightarrow \text{Heis}^3$ timelike horizontal geodesic lines issuing from zero 0, then satisfy to the following equations:

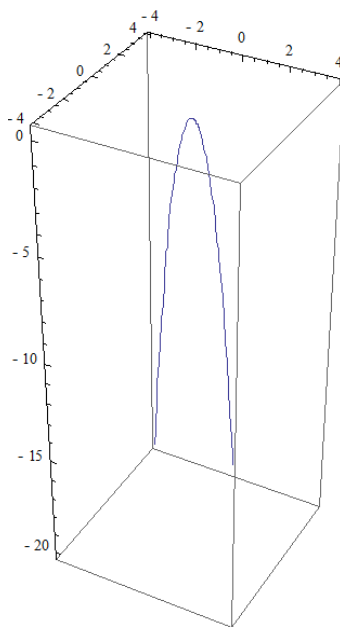
$$(3.13) \quad \begin{aligned} x(s) &= s \cosh \phi + a, \\ y(s) &= s \sinh \phi + b, \\ t(s) &= \frac{s^2}{2} \sinh \phi \cosh \phi + a \sinh \phi s + c, \end{aligned}$$

where a, b, c are constants of integration.

Next, we apply Corollary 3.9.

Example 3.10. Let us consider parametric equation of timelike geodesic lines with $a = b = c = 1$. Then, we can draw timelike horizontal geodesic lines with

helping the programme of Mathematica as follow



Theorem 3.11. *There exists no null horizontal geodesic in Lorentzian $Heis^3$.*

Proof. We prove that by contradiction. Let us assume that $c : I \rightarrow Heis^3$ is null horizontal geodesic curve parametrized by arc length. Then, null geodesic lines issuing from 0 in the Lorentzian $Heis^3$ is solution following differential equation system;

$$(3.14) \quad \begin{aligned} x'(s) &= |\xi| \sinh(2\xi s + \phi) \\ y'(s) &= |\xi| \cosh(2\xi s + \phi) \\ t'(s) &= \xi - x(s) |\xi| \cosh(2\xi s + \phi). \end{aligned}$$

Substitute equations (3.11) into (3.16) we have

$$(3.15) \quad \eta(s) = 0, \quad \rho(s) = 0.$$

From (3.8) we get $c'(s) = 0$. This contradiction proves the claim.

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